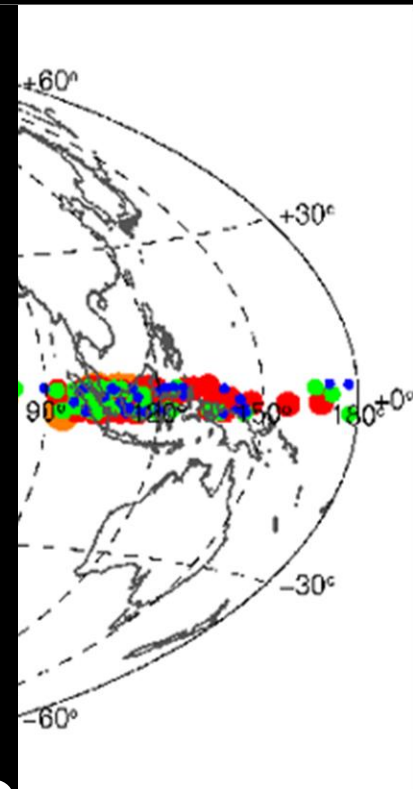
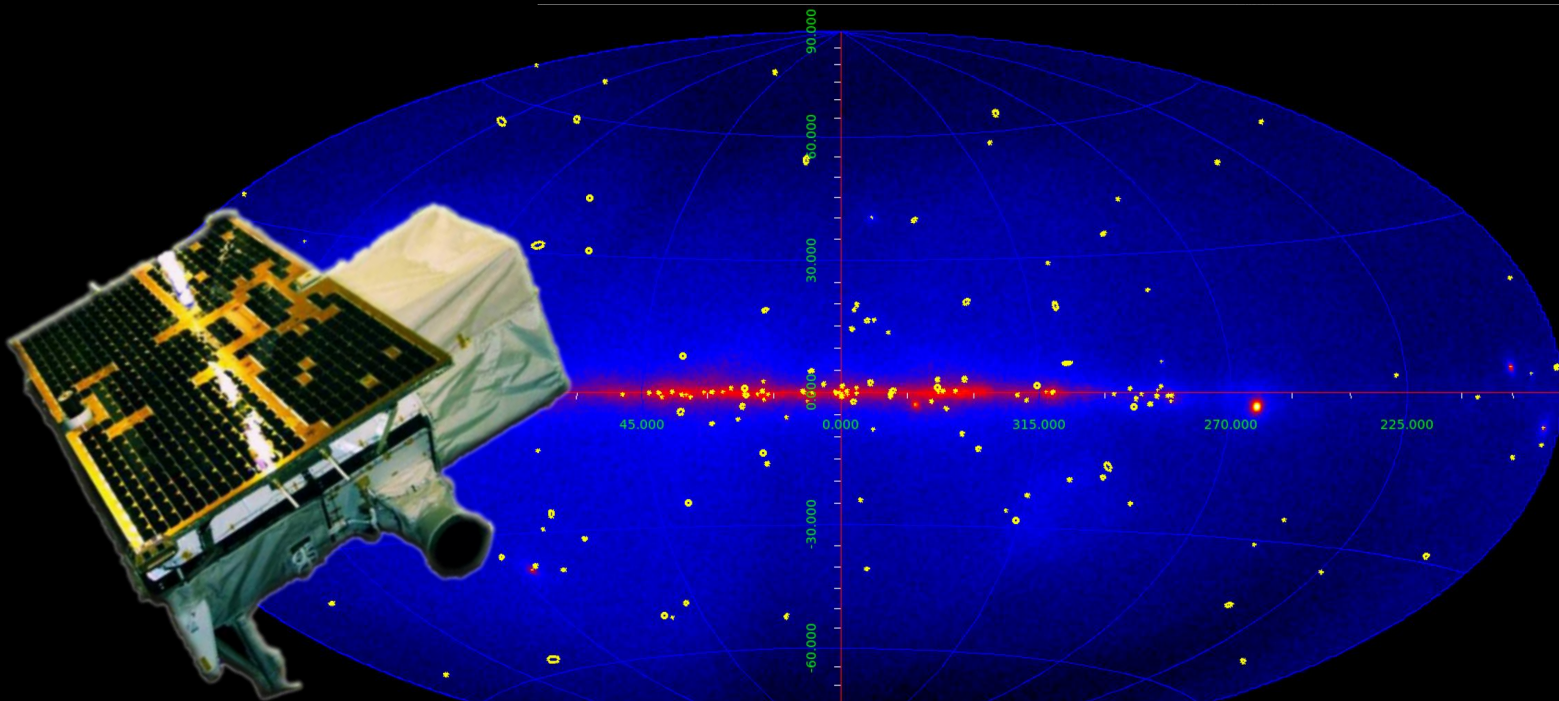




Overview of AGILE Catalogues, GRID & MCAL



**Francesco Verrecchia and Fabrizio Lucarelli
on behalf of the AGILE Data Center, ASI SSSC
& INAF-OAR**

Treasures Hidden in HE Catalogues, Toulouse, May 18, 2018

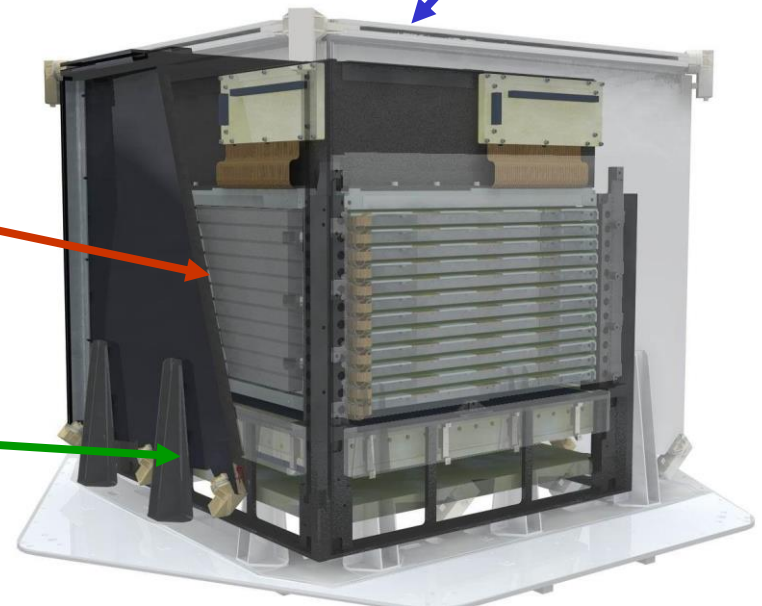
The AGILE instrument

- **AGILE is unique combination of X-ray and gamma-ray detectors ->GW counterpart search**
- **two co-aligned imaging detectors in hard X-rays (Super-AGILE) and gamma (GRID)
+ omnidirectional MCAL**

Super-AGILE
Coded aperture
18 - 60 keV
~1 sr FoV

Gamma Ray Imaging
Detector Silicon tracking
detector 30 MeV - 50 GeV
~2.5 sr FoV

Mini-calorimeter
non imaging scintillator
0.3 - 200 MeV
almost all-Sky FoV



The AGILE instrument

Gamma-ray imaging detector (GRID)

Energy range 30 MeV–50 GeV

Field of view ~ 2.5 sr

Flux sensitivity ($E > 100$ MeV, 5σ in 10^6 s) 3×10^{-7} (ph cm $^{-2}$ s $^{-1}$)

Angular resolution at 100 MeV (68% cont. radius) 4.3 degrees

Angular resolution at 1 GeV (68% cont. radius) 0.7 degrees

Source location accuracy (high Gal. lat., 90% C.L.) ~ 15 arcmin

Energy resolution (at 400 MeV) $\Delta E/E \sim 1 \Rightarrow$ +energy dispersion

In flight Crab radial profile from AGILE Pointing mode data compared to the Fermi/LAT «front» event one (P7REP).

Sabatini et al. 2015

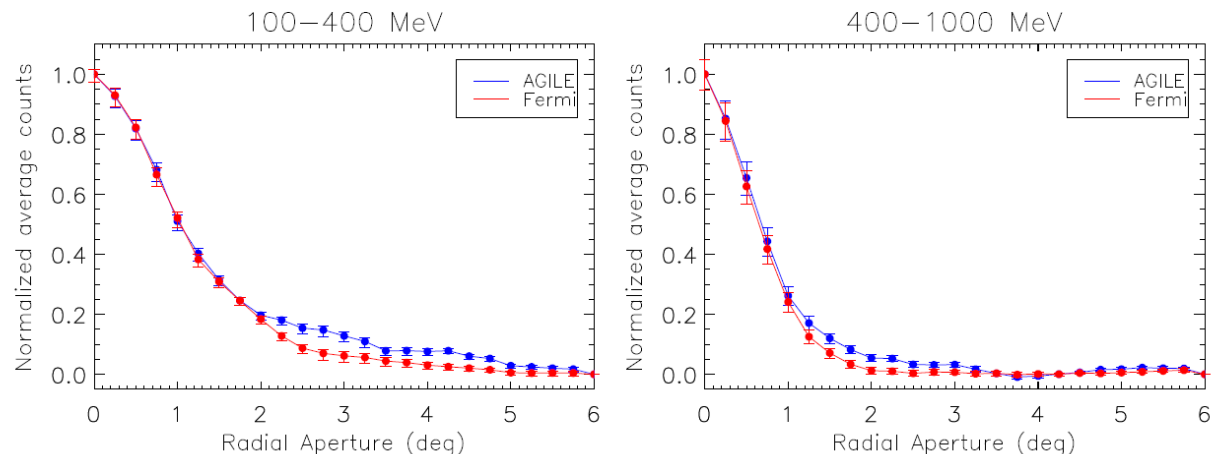


Figure 8. Average count radial profiles for circular apertures of increasing radii at steps of 0:25 of the Crab (pulsar + Nebula); in-flight data for the AGILE/GRID (blue data points) and Fermi/LAT (red data points). Left panel: 100–400 MeV energy range. Right panel: 400–1000 MeV energy range.

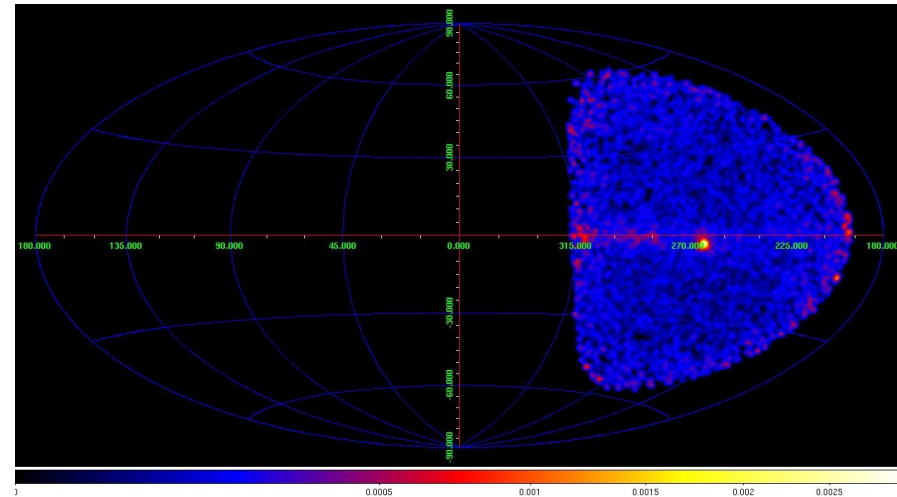
\Rightarrow photon-by-photon data

The "two" AGILE lives

- **Pointing mode till Nov 2009, 60° radius FoV**

coverage: 1/5 of the sky

2d exposure

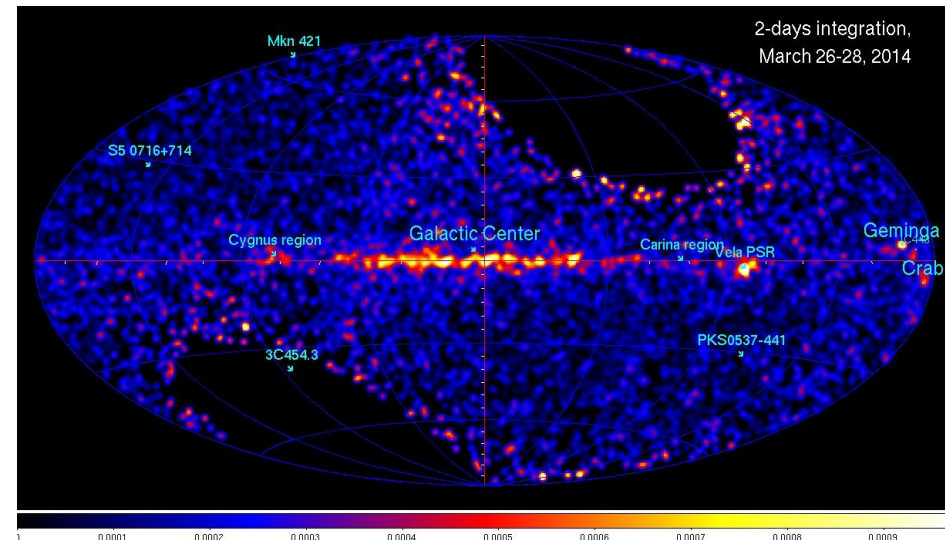


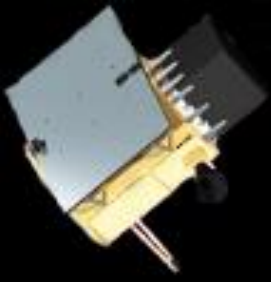
- **Spinning mode after, sky scanning (with solar panels constraints)**

coverage: about 80% of the sky in

~ 7 min

2d exposure





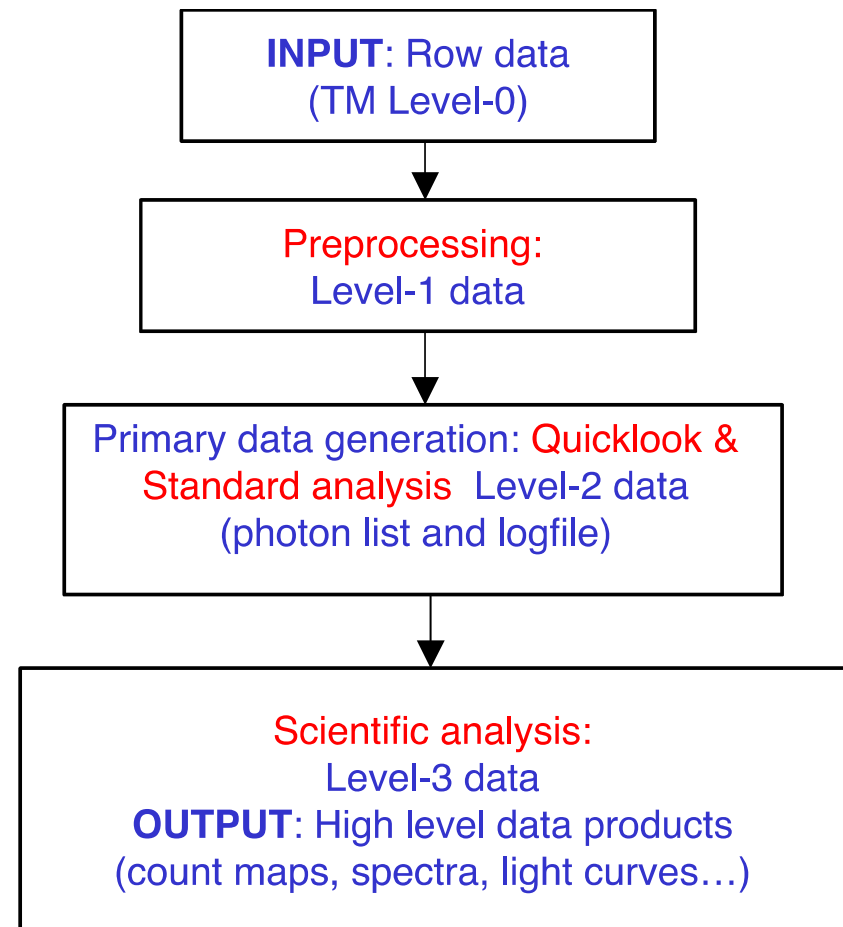
AGILE

Science Data Center

- The ADC, based at SSCDC c/o ASI-HQ (Rome), is in charge of **all the scientific oriented activities related to the analysis and archiving** of AGILE data:

the processing and archiving of all the data levels, from scientific telemetry (TM) Level-0,

- Preprocessing ->Level-1
- Creation of main archive -> Level-2 (photon lists and spacecraft files), within Quick-Look & Standard analysis
- Scientific analysis -> Level-3 scientific maps (counts, exposure and diffuse gamma-ray background maps), for various procedures ->source detections (positions, flux and significances), spectra, light curves



AGILE Science Alert System

- The system is distributed among the ADC @ SSDC and the AGILE Team Institutes (Trifoglio, Bulgarelli, Gianotti et al.)
- Automatic Alerts to the AGILE Team are generated within $T_0 + 100$ min (GRID)
- GRID Alerts are sent via email (and sms) both on a contact-by-contact basis and on a daily timescale
- Refined manual analysis on most interesting alerts performed every day (daily monitoring) by dedicated advocates

Standard procedure for AGILE-GRID gamma-ray “all data” catalogues:

catalog of sources on incremental exposure time

0) LV2 archive update: event list creation =>AGILE “event filter” based on a Kalman filter
→select event of different class types having decreasing probability to be a photon

a) Level-3 scientific products: creation of scientific maps (counts, exposure and diffuse gamma-ray background maps) on the “whole dataset”, **integrated**, centered at **HEALPix** pixels, in standard full band **$E > 100 \text{ MeV}$** , and in a set of sub-bands to later evaluate source spectra

b) Preliminary blind search algorithmS: “source candidate positions” (feeds) extraction, AGILE “spotfinder”, Ximage/detect, SExtractor -> candidate source list

c) scientific analysis: AGILE “multi-source” Maximum Likelihood Estimator (MLE) analysis, evaluating Test Statistics (TS) build as the ratio of the ML functions for the null (background only) and the alternate (source presence) hypotheses. Evaluate source parameters (position, significance, flux, spectral index and errors) and diffuse emission parameters =>**iterative algorithm, but due to high number of sources & to reduce the parameter space, we execute an iterative procedure:**

- **First step:** evaluate first guess for source fluxes and then SINGLE source position refinement and galactic diffuse parameters, keeping fixed a mean spectral index (-2)
- **2nd step:** repeat flux and significance evaluation at fixed position&diffuse par. ->reiterate (?)
- **Final:** estimation of spectra for most significant sources

Concluding source selection based on significance and MLE 95% c.l position contours

Other types of catalogues

II: AGILE-GRID gamma-ray “all data” catalogues of preselected class of sources

In this case source positions are known in advance -> mainly execution of MLE **at fixed position** with a **lower significance threshold** ($\geq 3\sigma$) and/or allowing positions refinement for specific class of sources

=>For Pointed observations:

III: AGILE-GRID gamma-ray “single observations”:

Similar procedure as first type of catalog but on usually a single or few maps; known dependence of source detection with the off-axis angle with respect of the map center

Review of AGILE-GRID catalogues on long integration time scales.....

.....not exhaustive:

- all on Pointing mode**
- not considering transients (days)**

Treasures in a complete GRID archive still not completely investigated

All catalogs on-line version from ADC web page <http://agile.ssdsc.asi.it/>

The First AGILE GRID Catalogue of γ -ray Sources

Period July 2007 -- June 2008

Pulsars

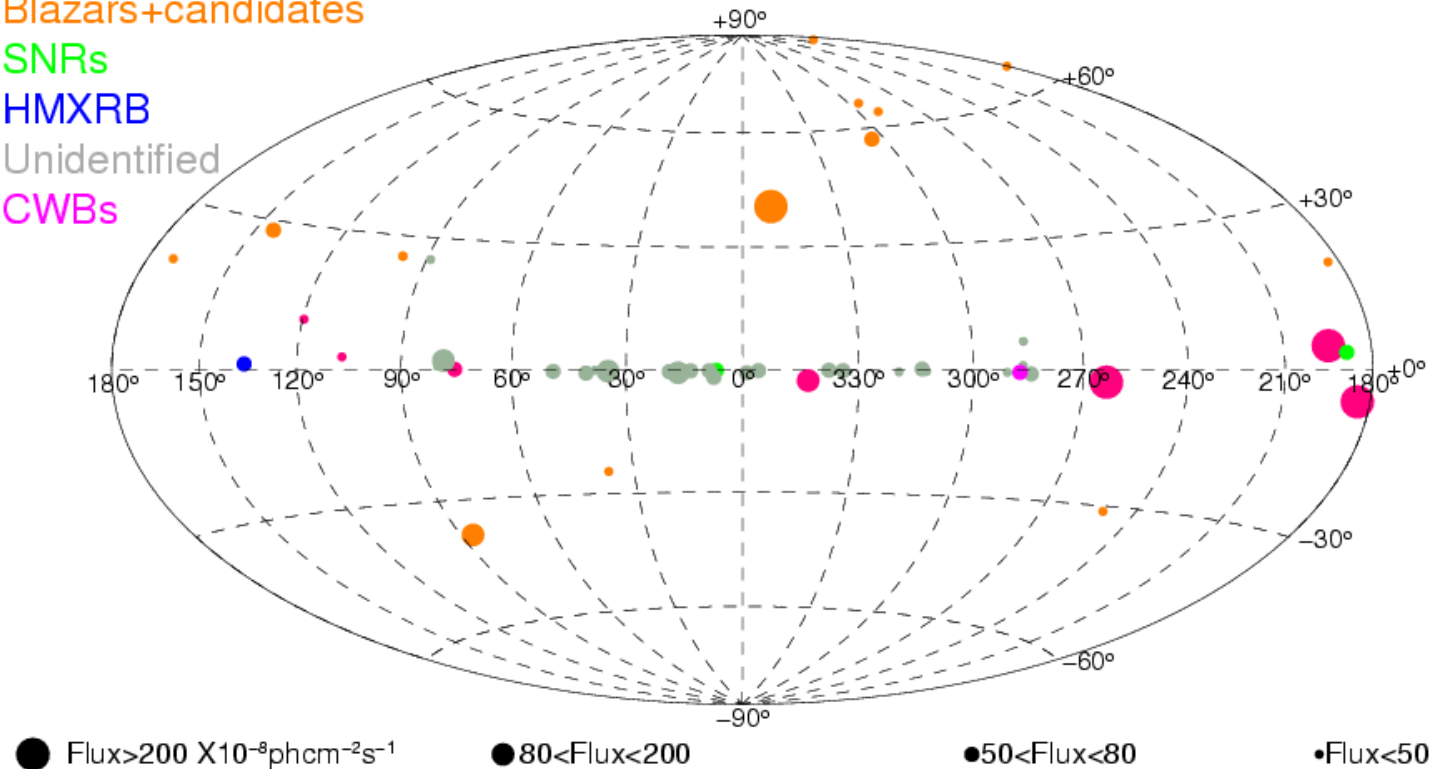
Blazars+candidates

SNRs

HMXRB

Unidentified

CWBs



C. Pittori et al., A&A 506, 2009

47 high confidence 1AGL sources $E > 100 \text{ MeV}$:

- 21 confirmed and candidate **Pulsars**,
- 13 **Blazars** (7FSRQ, 4BL Lacs, 2 unknown type),
- 2 possible **HMXRBs**,
- 2 possible **SNRs**,
- 1 **Colliding-wind Binary System** (Eta-Car)
- 8 «**Unidentified**» sources.

An updated list of AGILE bright γ -ray sources and their variability in pointing mode

(F. Verrecchia et al., 2013, A&A, 558, A137)

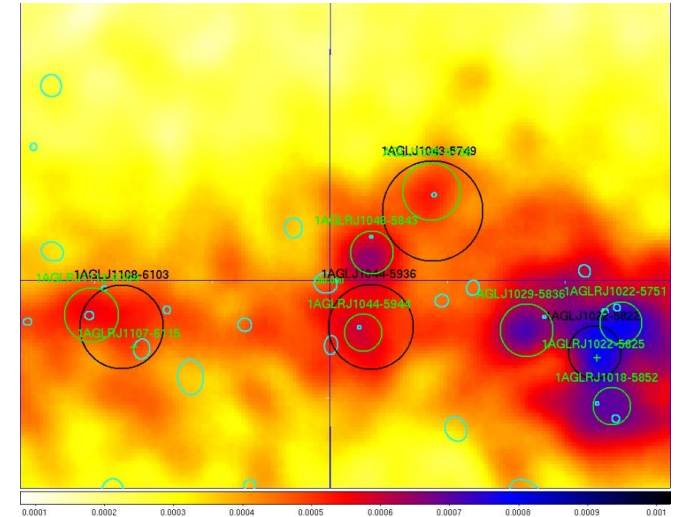
Variability study of an **improved source list** compared to 1AGL, preliminarily revised on the 2.3 yrs “all data” maps of the whole AGILE pointed observations (OB)

No complete source blind search executed, similar to type II catalogues

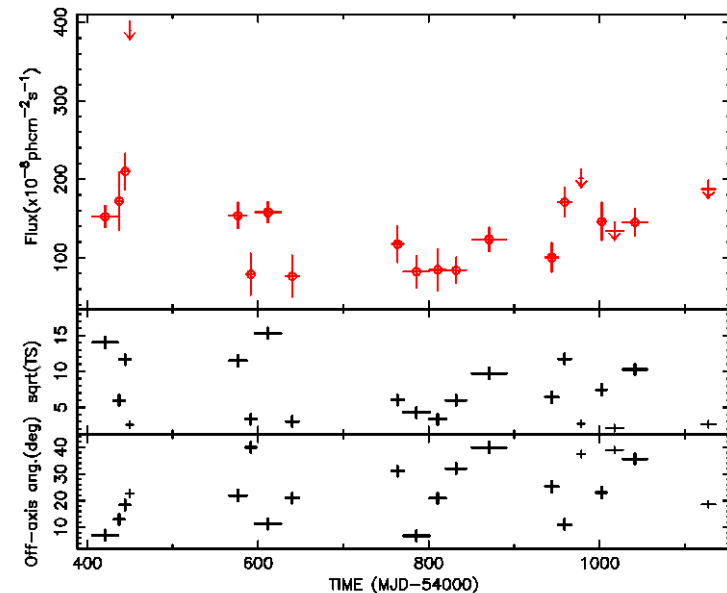
Processing of EACH SINGLE Obs (duration 1-45d) in $E > 100$ MeV band allowed to detect 54 sources, among which 15 new one =>variability results

OB time scale
light curves \rightarrow

Refined positioning of some 1AGL sources: the Carina/W2 region



1AGLR J2021+4030, PSR/SNR GammaCygni, OB Processing

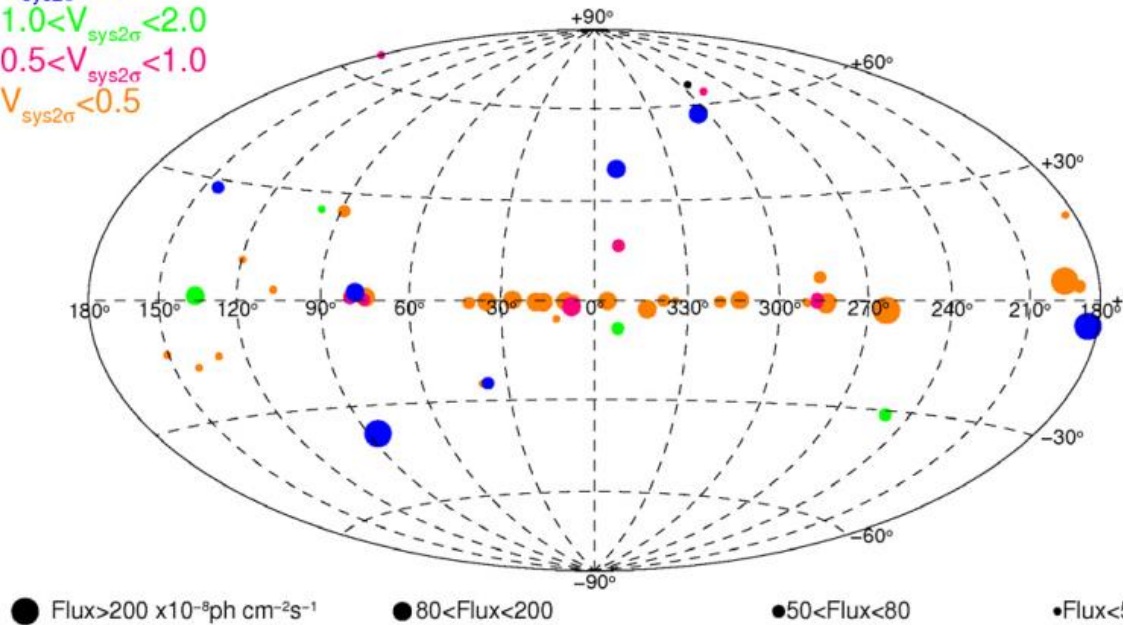


An updated list of AGILE bright γ -ray sources and their variability in pointing mode

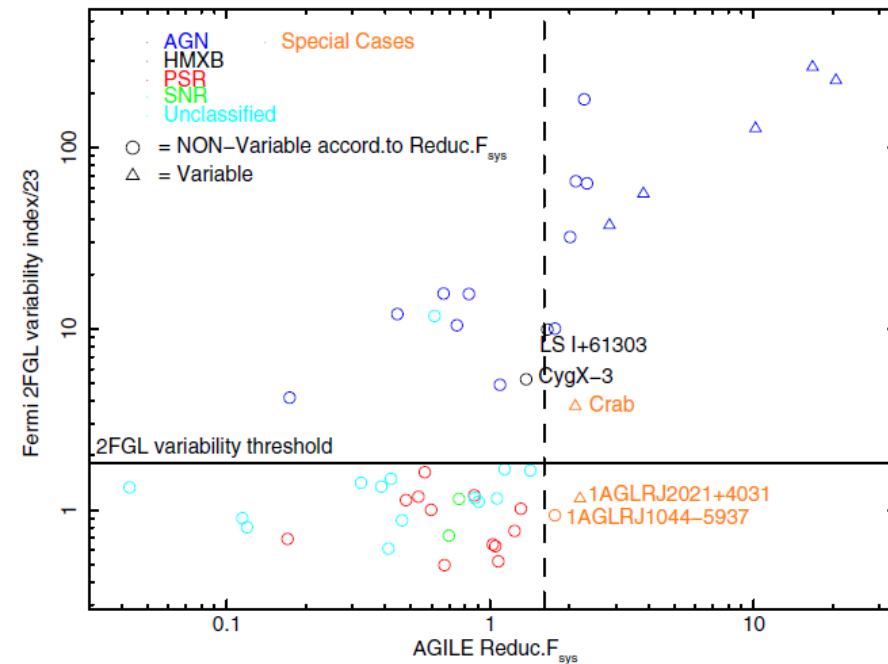
Variability indices compared with 1FGL, 2FGL: 12 well known variable sources, named **1AGLR**

Single detections

$V_{\text{sys}2\sigma} > 2.0$
 $1.0 < V_{\text{sys}2\sigma} < 2.0$
 $0.5 < V_{\text{sys}2\sigma} < 1.0$
 $V_{\text{sys}2\sigma} < 0.5$



Comparison with 2FGL variability index (assoc.radius=30')



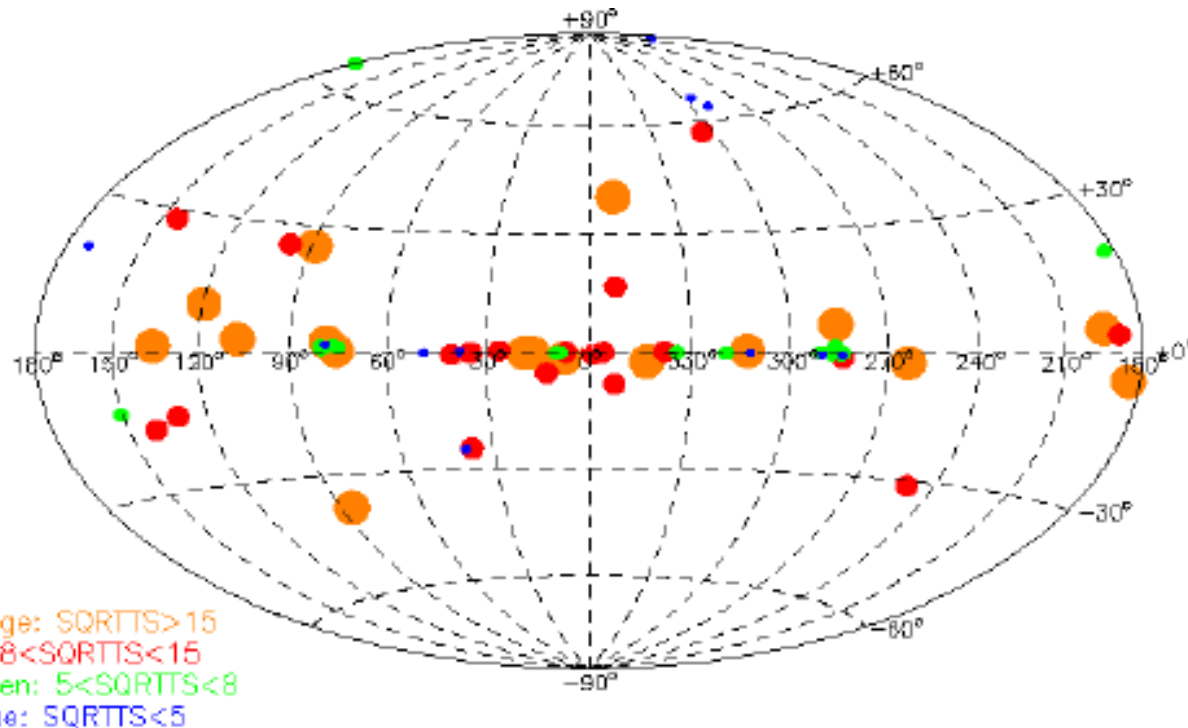
No mean fluxes on “ALL data” maps published.

The AGL-ALL catalogue project: incremental catalogue

The 1AGLR lack of updated **mean fluxes** on all pointing data => new calibration delivered so need to have homogeneous “mean” source parameters.

Executed a data processing in $E > 100$ MeV band on an input list of 65 1AGL + 1AGLR sources on new Pointing dataset maps generated with new calibration

=>defined a new procedure at fixed source position (cat of type II)



Results:

61/65 significant 1AGL/1AGLR,

4 sources not detected,

=>flaring sources or too weak

Source list presented at 2015 13th AGILE Workshop

Then updated with new sources from specific publications

Search of MeV-GeV counterparts of TeV sources with AGILE in pointing mode

A. Rappoldi, F. Lucarelli, C. Pittori, F. Longo, P.W. Cattaneo, F. Verrecchia, et al., A&A, 2016

Exploring TeV-GeV connection






TeV sources input list: 147 TeV source positions taken from the **TeGeVCat** (Carosi et al. 2016 ->previous Lucarelli's talk)

Procedure: type III catalogue

- dataset: the **pointing mode** observations (Jul 2007—Oct 2009)
- Input list of TeV source + AGL-ALL reference list
- Iterative analysis procedure based on MLE with source detection and
- localization in two steps
 - **fixed** position (at the TeVCat coordinates)
 - **free** position (near to the original one)
- **Spectral** analysis for the most significant ones

for TeV source counterparts search possible displacement was investigated, comparing it with TeV source “extension” and poor angular resolution

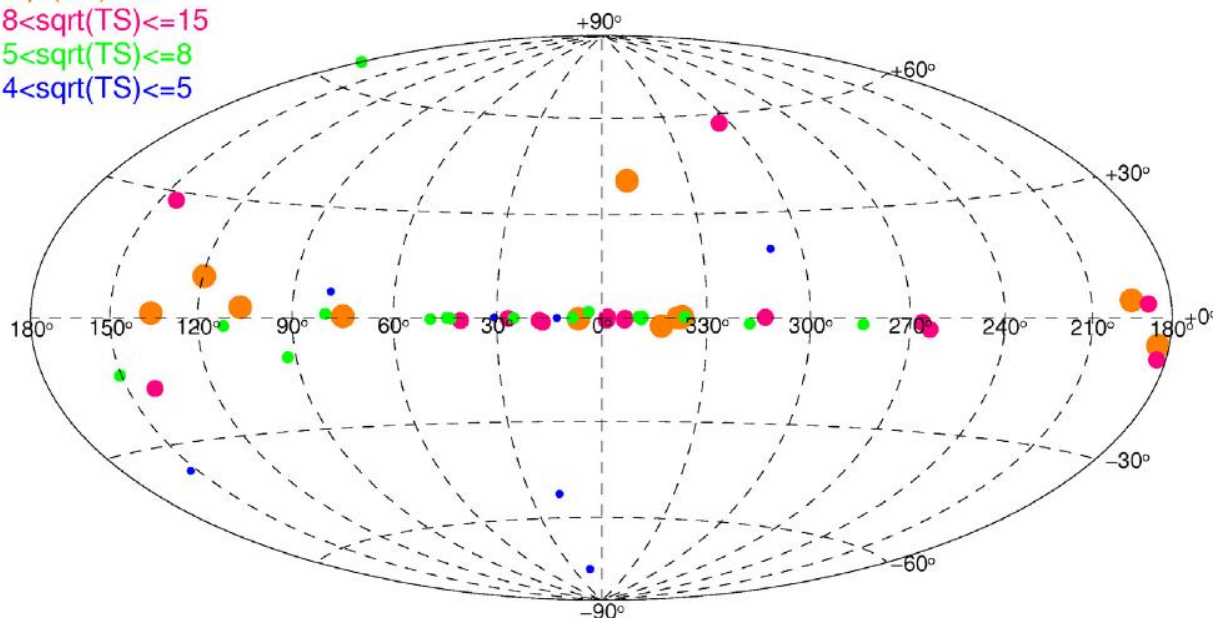
Source Types

-  PWN
-  Binary XRB PSR Gamma BIN
-  HBL IBL FRI FSRQ LBL AGN (unknown type)
-  Shell SNR/Molec. Cloud Composite SNR
-  Starburst
-  DARK UNID Other
-  uQuasar Star Forming Region Globular Cluster Cat. Var. Massive Star Cluster BIN BL Lac (class unclear) WR

Results: known and new sources

- **52** TeV sources show a significant *count excess* in the **AGILE** data in pointed observations period, 35% of the original sample
- **26** have a spatial association with already known **AGILE** sources from 1AGL/1AGLR catalogs: **15** galactic, **6** extra-galactic, **5** unassociated
- **26** detections represent new **AGILE** sources (with respect to the reference catalogs): **15** galactic, **7** extra-galactic, **4** unidentified

- $\sqrt{TS} > 15$
- $8 < \sqrt{TS} \leq 15$
- $5 < \sqrt{TS} \leq 8$
- $4 < \sqrt{TS} \leq 5$



Source Type	Detected / Total	Source Class	Detected / Total
Extra-galactic	13 / 61 (21%)	Blazar	0 / 1 (0%)
		HBL	5 / 44 (11%)
		IBL	2 / 5 (40%)
		LBL	2 / 3 (67%)
		FSRQ	2 / 3 (67%)
		Sbs	0 / 2 (0%)
		FRI	2 / 3 (67%)
Galactic	30 / 58 (52%)	PWN	11 / 28 (39%)
		SNR	7 / 11 (64%)
		PWN/SNR	2 / 2 (100%)
		SNR/MC	5 / 8 (63%)
		BIN/XRB	3 / 5 (60%)
		GC	1 / 1 (100%)
		WR	1 / 3 (33%)
Unidentified	9 / 28 (32%)	---	---

THE SECOND AGILE CATALOG OF GAMMA-RAY SOURCES: THE FINAL REVIEW

A. BULGARELLI, N. PARMIGGIANI, V. FIORETTI, M. TAVANI, G. PIANO, M. CARDILLO
(INAF)

C. PITTORI, F. VERRECCHIA, F. LUCARELLI (SSDC AND INAF)

A. RAPPOLDI, P. CATTANEO (INFN-PV)

A. ABOUDAN (CISAS)

- The 2th AGILE Catalog of gamma-ray sources: the final review

Evolved a lot through the last years following various calibration versions and sw improvements.

After development of new BUILD sw (SCI) and new calibration files (IRFS)

- AGILE/GRID observations covering the time period July 4, 2007, to October 15, 2009 (the AGILE POINTING MODE).
- The analysis is based on data in the
 - 100 MeV to 50 GeV energy range
 - A check in the "low" energy range 30 MeV - 100 MeV has been performed for the most significant sources
- Source detection is based on the integrated data set, i.e., sources are detected according to their **average fluxes** over about 27 months.

UPDATES

- A long path:
 - The catalog adopts now different spectral shapes
 - Simple Power Law
 - Power Law with exponential cutoff
 - Power Law with super exponential cutoff
 - Log Parabola
 - The NEW BUILD25 (AGILE/GRID Science Tools) has been used, that include also the **energy dispersion correction** factor (EDP)
 - New Instrument Response Functions: H0025 -> energy range
- ➔ **spectral comparison** with Fermi/LAT: same spectral bins till 10GeV

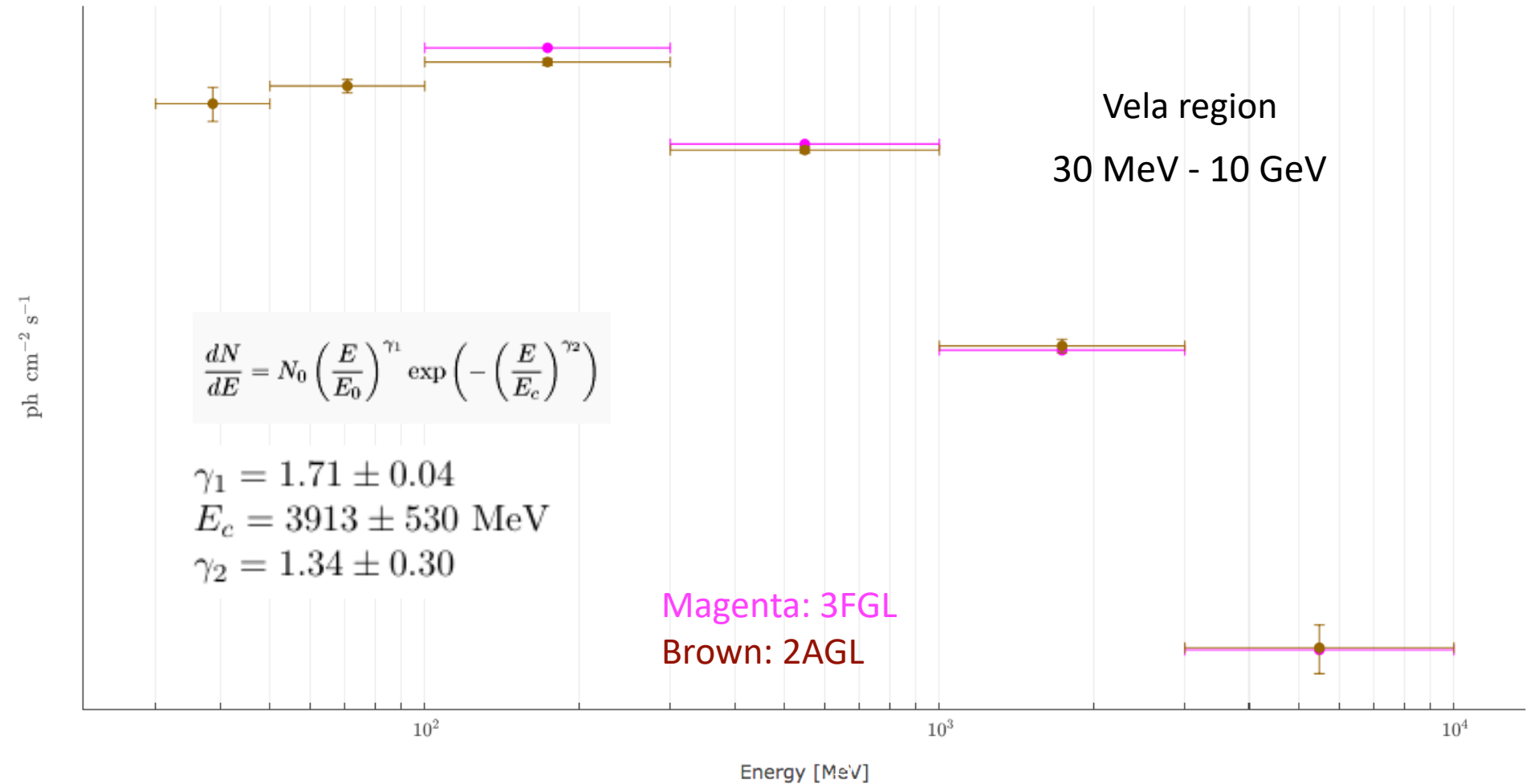
$$\frac{dN}{dE} = N_0 \left(\frac{E}{E_0} \right)^\gamma$$

$$\frac{dN}{dE} = N_0 \left(\frac{E}{E_0} \right)^{\gamma_1} \exp \left(- \left(\frac{E}{E_c} \right) \right)$$

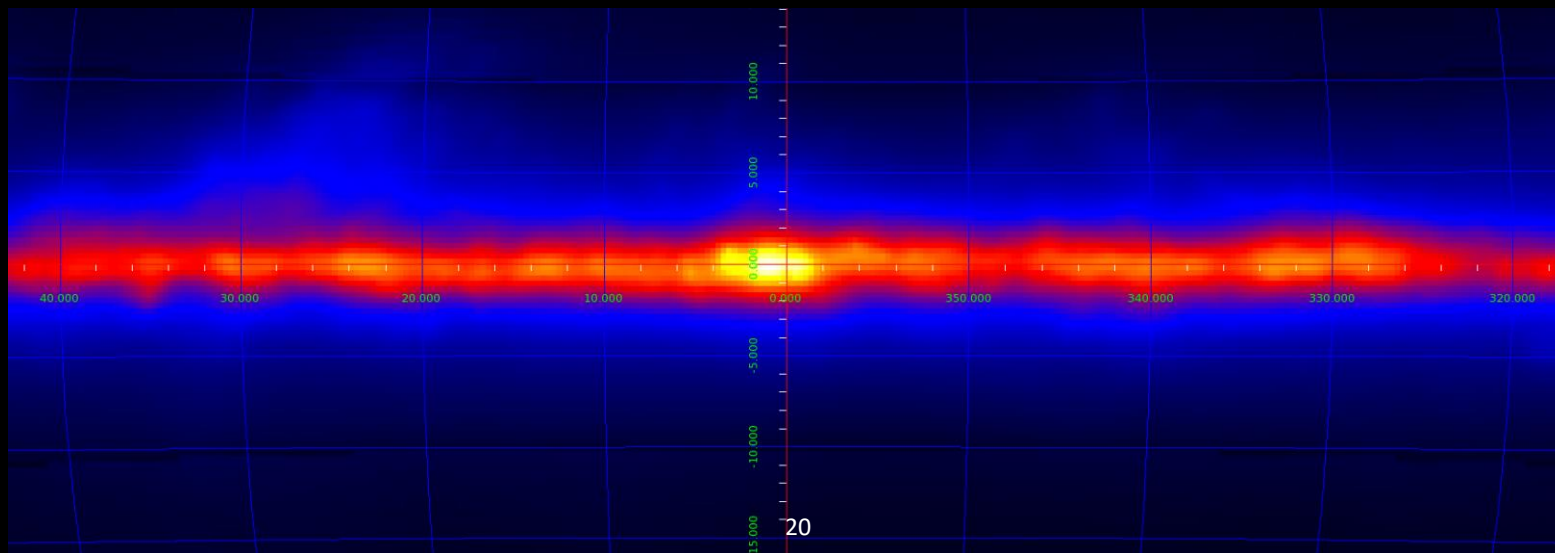
$$\frac{dN}{dE} = N_0 \left(\frac{E}{E_0} \right)^{\gamma_1} \exp \left(- \left(\frac{E}{E_c} \right)^{\gamma_2} \right)$$

$$\frac{dN}{dE} = N_0 \left(\frac{E}{E_b} \right)^{-(\alpha + \beta \log(E/E_b))}$$

NEW BUILD25: CURVED SPECTRAL ANALYSIS



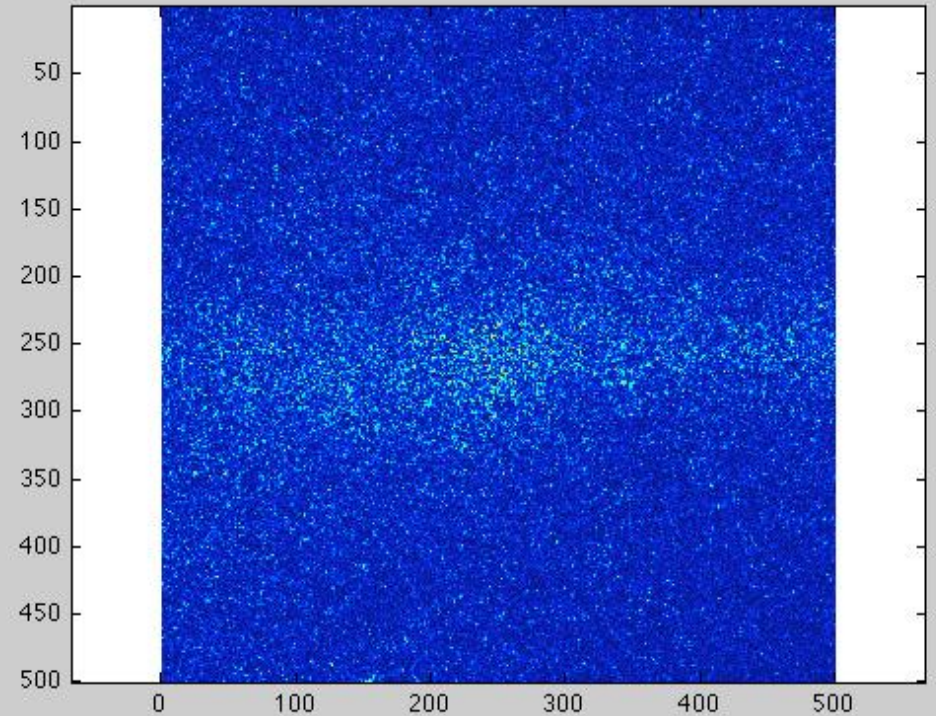
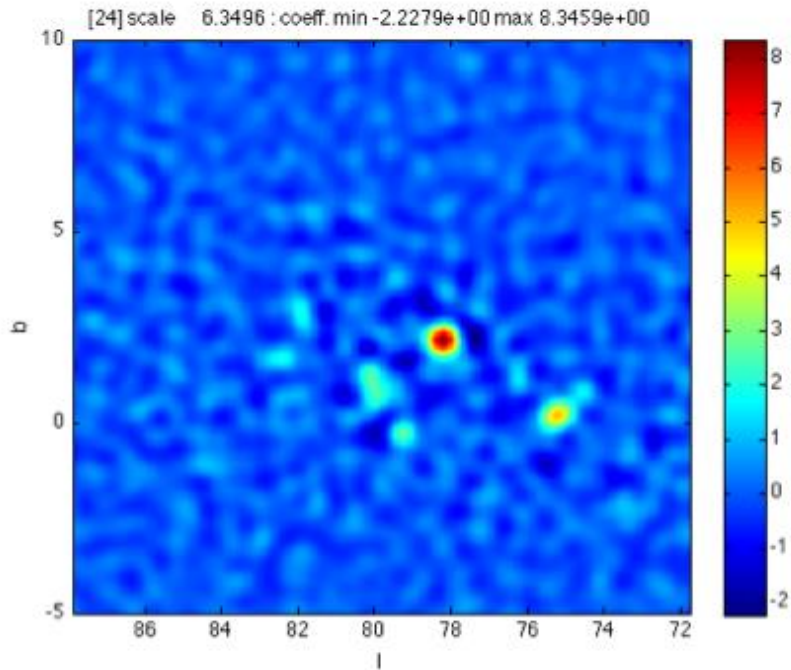
- Cross check with previous BUILD22@SSDC: only confirmed detections are included in the catalog
- The gamma-ray diffuse model used in the analysis has been improved. The diffuse model is particularly important for sources from low to mid Galactic latitudes.



First step: the seeds

source blind search

WAVELET ALGORITHM (THE SEED METHOD N° 1)



Second step

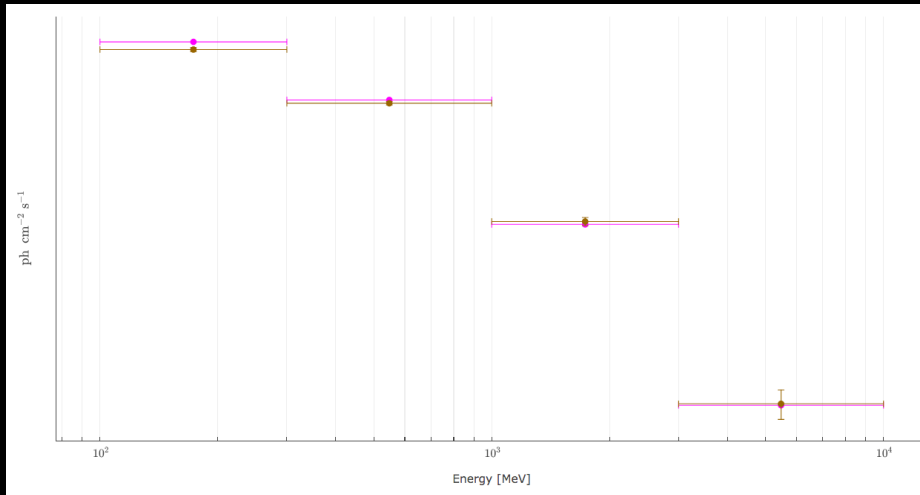
MANUAL ANALYSIS

- From the list of seeds using all methods...
 - Manual analysis
 - Final determination of position and flux

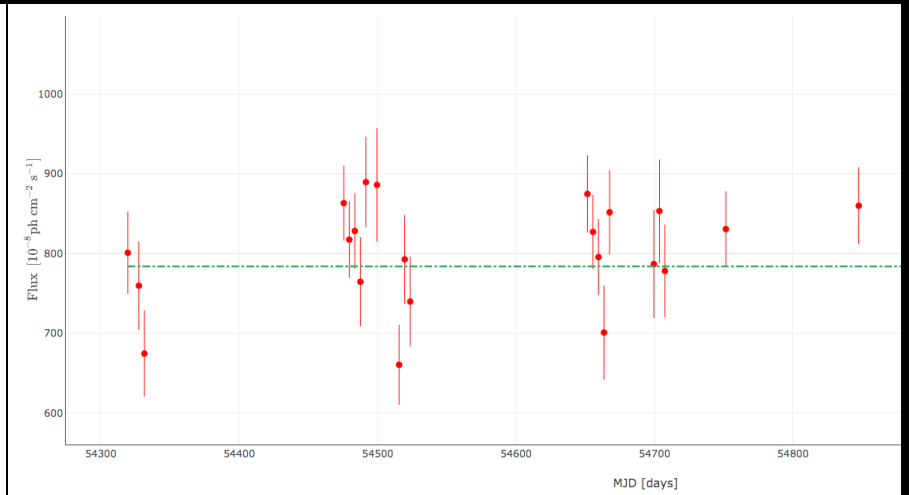
Third step

CHECK THE RESIDUAL TS MAP

PRODUCTS



Spectra (100-10000 MeV)



Light curves (4 days)

Lets TS_0 the value of TS evaluating all the time bins at the same time but considering a constant flux, TS_1^i the value of TS optimizing the flux in each period of time i .

The variability index is

$$TS_{var} = \sum_{i=1}^N TS_1^i - TS_0$$

If the null hypothesis is correct TS_{var} is distributed as χ^2 with N degrees of freedom, and a value of $TS_{var} > h(N)$ is used to identify variable sources at a 99% confidence level.

It is possible to introduce a corrective factor (similar to (Nolan 2012)):

$$TS_i^{corr} = F_{sigma_i}^2 / (F_{sigma_i}^2 + f^2 * F_0^2)$$

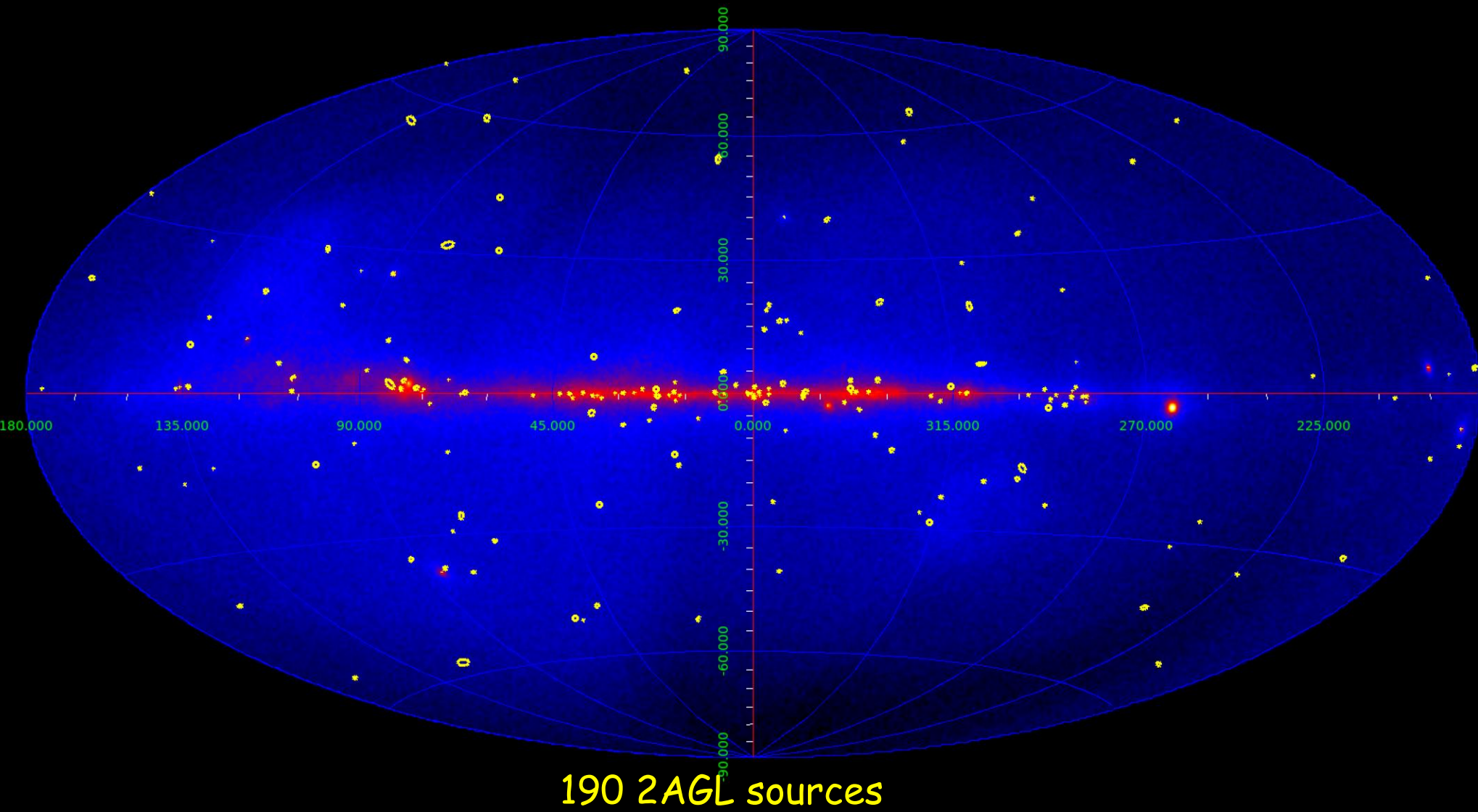
We consider $f = 0.01$ in our analysis. The variability index is

$$TS_{var}^* = \sum_{i=1}^N (TS_i^{corr} * TS_1^i) - TS_0$$

Curvature Index

Variability Index

RESULTS



PRELIMINARY ASSOCIATIONS

- Positional coincidence only
- These associations are not in general to be taken as firm identifications: a physical relationship²⁵ is not established between gamma-ray sources and sources in other wavelengths.
- 20 AGILE-only (no 3FGL) sources

AGN

Description	number
BCU (Blazar candidate of uncertain type)	10
BLL (BL Lac type of blazar)	17
FSRQ (FSRQ type of blazar)	31
RDG (Radio galaxy)	2 (CenA, NGC1275)
BIN (Binary)	1 (Eta Carinae)
GLC (Globular cluster)	1 (Terzan 5)
HMB (High-mass binary)	3
PSR (positional only)	50
PWN	3
SNR	9
SPP	7
	134

HMB (High-mass binary)	LSI+61 303, 1FGLJ1018.6-5856, Cygnus X-3
SNR	IC443, CTB37A, W28, W30, W44, W49B, W51, GammaCygni, HB21
PWN	Crab Nebula, HESSJ1632-478, PWNG0.13-0.11)

FUTURE:

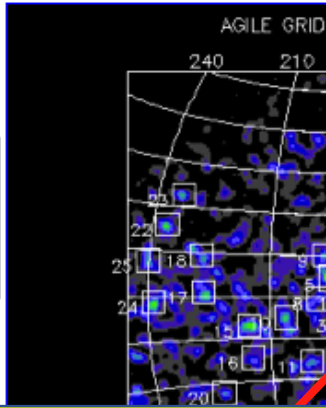
- New catalogue «integrated» maps: including Spinning mode data
→ lower acquisition efficiency, currently used in specific work
- New homogeneous «transient» source catalogue: on both archives ; Variability catalog(s) based on experience from 2AGL and using ADC light curve tool (**ALCE**) and LV3 tool (Lucarelli's talk)

SSDC Data Explorer

Quick Look AGILE data

Available parameters

- Name
- Ra Dec
- Gal Iso
- Cnts Cnts
- Err. Err.
- Sqrt(TS)
- XimageId XimageId
- Flux Flux
- Err. Err.
- Distance from FOVCent.
- Ximage SNR
- Sp_Index
- Err_sp_index
- Other_name



Example of the ALCE Tool, simple version

The ASDC Multi-frequency Data Explorer: Web and VO ...

wwwdev.asdc.asi.it/showEntry.php#

ASDC All Science Data Center Version 2.0

Entry 1AGLJ0535+2205 --- Crab*
 R.A. (J2000) = 05 35 06.0 (83.7750 deg) l=184.56
 Dec (J2000) = +22 05 41.7 (22.0949 deg) b=-5.63
 Galactic nH = 3.27E+21 (cm^-2) Source Names

AGILEPUB GRID data products Error circle EXPLORER Source Details Feedback Names

Catalog: aglIvar Radius: 1 Start Time: End Time:
 (Select DATE to loop the range in DD-MN-YY format or MJD to loop)

Duration (days): Sample: FM Sqrt(TS) >: Additional Y plot:

1AGL10535+2205 LC All Data Table
 1AGL10535+2205 LC Flux-only Data qdp

1AGLJ0535+2205

FLUX 10u-8d ph cmu-2d au-1

5.47x10⁴ 5.48x10⁴ 5.49x10⁴ 5.5x10⁴

TIME (MJD)

alce_dacf_2FGLJ0210.7-5102_T_qdp

DACF

Time lag [days]

DACF data QDP
 DACF data

Time_Lag[days] DACF err_DACF

ZDCF PARAMETERS:
 Autocorrelation? T
 Uniform sampling? F
 Omit zero lags? F
 Minimal # in bin: 11
 # of MonteCarlo: 100
 MonteCarlo seed: 123
 132 points in /web/wwwdev/wwwscr/30438qdp.qdp.dat

Binning with minimum of 11 points per bin and resolution of 1.00E-03 .
 163 bins actually used, 4915 inter-dependent pairs discarded.

tau	sig(tau)	-sig(tau)	dcf	err(dcf)	-err(dcf)	(#bin)
0.000E+00	0.000E+00	0.000E+00	1.000E+00	5.960E-08	-5.960E-08	(132)
2.000E+00	3.060E+00	-1.092E+00	3.904E-01	9.840E-02	-1.031E-01	(73)
1.000E+00	5.404E+00	-1.764E+00	4.044E-01	8.784E-02	-9.178E-02	(89)
2.110E+00	7.733E+00	-1.933E+00	3.860E-01	9.476E-02	-1.031E-01	(76)
3.119E+00	1.407E+00	-9.211E-01	3.233E-01	1.173E-01	-1.226E-01	(58)
3.883E+01	5.263E+00	-1.744E+00	2.688E-01	1.109E-01	-1.146E-01	(70)
5.067E+01	5.785E+00	-2.743E+00	3.452E-01	1.033E-01	-1.077E-01	(72)
6.324E+01	6.252E+00	-2.822E+00	2.610E-01	1.157E-01	-1.197E-01	(65)
7.685E+01	5.840E+00	-2.820E+00	2.054E-01	1.277E-01	-1.313E-01	(57)
8.896E+01	4.783E+00	-2.365E+00	1.205E-01	1.414E-01	-1.440E-01	(50)
9.788E+01	3.537E+00	-1.287E+00	1.342E-02	1.672E-01	-1.676E-01	(38)
1.065E+02	3.850E+00	-2.360E+00	-5.274E-02	1.450E-01	-1.430E-01	(58)
1.180E+02	4.747E+00	-3.718E+00	2.841E-01	1.277E-01	-1.331E-01	(52)
1.301E+02	6.223E+00	-2.855E+00	3.528E-01	1.234E-01	-1.308E-01	(50)
1.407E+02	1.640E+00	-6.293E-01	1.130E-01	2.234E-01	-2.308E-01	(21)
1.454E+02	2.961E+00	-1.274E+00	1.640E-01	1.787E-01	-1.845E-01	(31)
1.523E+02	5.054E+00	-2.080E+00	1.176E-01	1.644E-01	-1.612E-01	(39)
1.777E+02	1.029E+00	-5.093E-01	1.509E-01	1.509E-01	-1.509E-01	(39)

Chi-squared/d.o.f. = 153.7/30 RA = 05 35 06.0
 Minimum value = 151.699997 Dec = 22 05 41.7
 Maximum value = 616.599976 Radius = 1
 Mean = 296.077423
 Weighted mean = 279.801806
 average = 296.077423
 median = 221.460009
 standard deviation = 98.7449546
 variance = 9750.56836
 skewness = 1.22759724
 kurtosis = 1.91491201

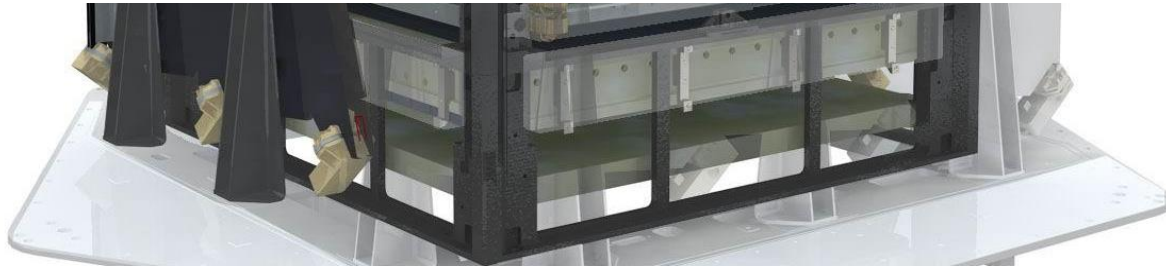
McLaughlin Variab. Index
 Sqrt(TS) > 3 Submit
 McLaughlin V (statist. errors + 10% systematic) = 3.418

Discrete AutoCorr. Est. (DACF)
 Minimal number of points per bin: 11
 Monte Carlo runs for error estimation: 100
 Submit

Entry number	AGILE name
1 Select	ASDC Data Explorer AGL J0634+1816
2 Select	ASDC Data Explorer AGL J1049+8055
3 Select	ASDC Data Explorer AGL J0832-1236

1-Apr-2012 20:35

MCAL catalogues



MCAL GRB Catalog

(M. Galli et al., 2013)

Photon-by-photon data from launch till Oct 2009: high-energy emission from GRB. →detected 85 GRB, 24 with spectral data

- Similar bimodal T_{90} distribution to lower energies, with 21% short GRB

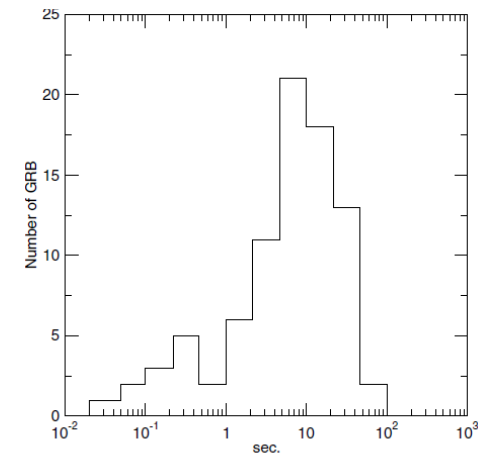
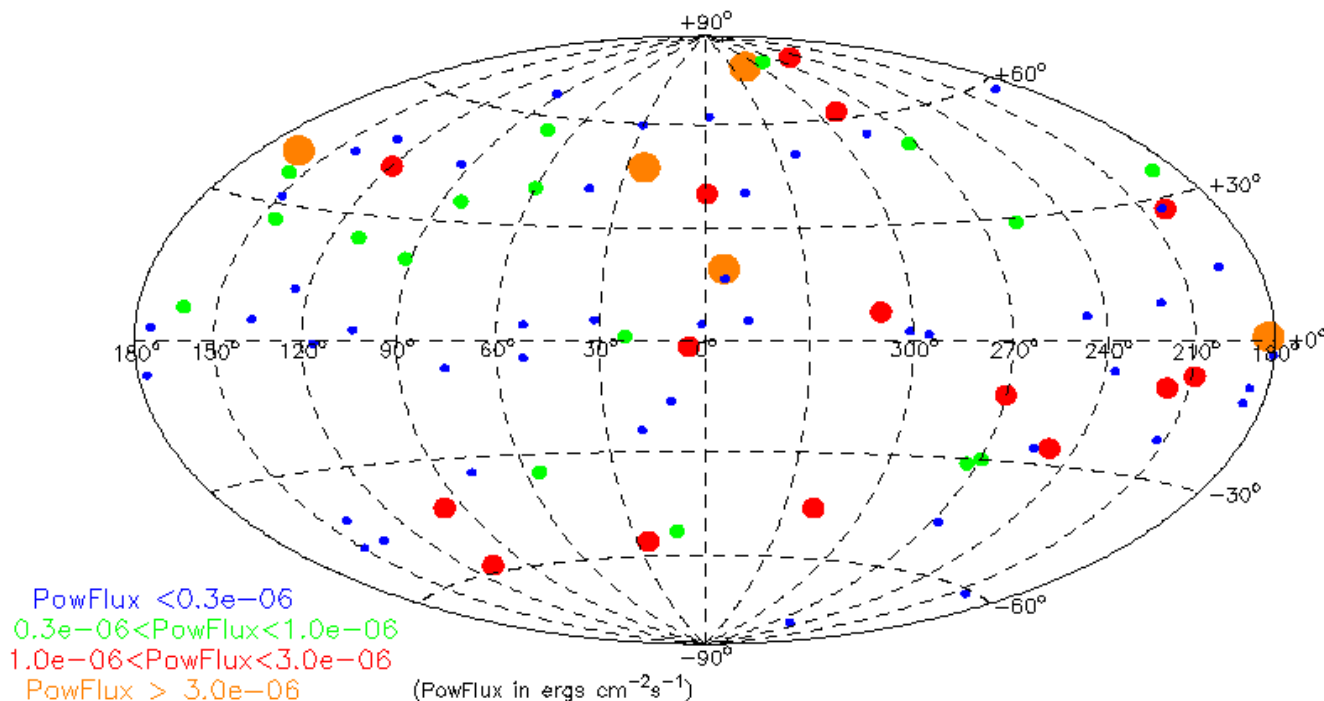


Fig. 3. T_{90} distribution.

→ **a new catalog** is in progress after development of new pipeline for GW counterpart searches

MCAL GRB Catalog (M. Galli et al., 2013) ADC interactive webpage

GRB observed from An

Entr
 (000) = 22 14 12
 (000) = -26 36 0
 nH = 1.66E+20

Swift-XRT light curves of GRB 090510

Last updated after receiving ObsID 00351588001, version 19

Related pages: [Burst Analyser](#) | [Enhanced position](#) | [Spectrum](#) | [GRB Region information](#) | [XRT Catalogue entry](#) | [Download obs data](#) | [GCN Notices](#) | [GCN Circulars](#)

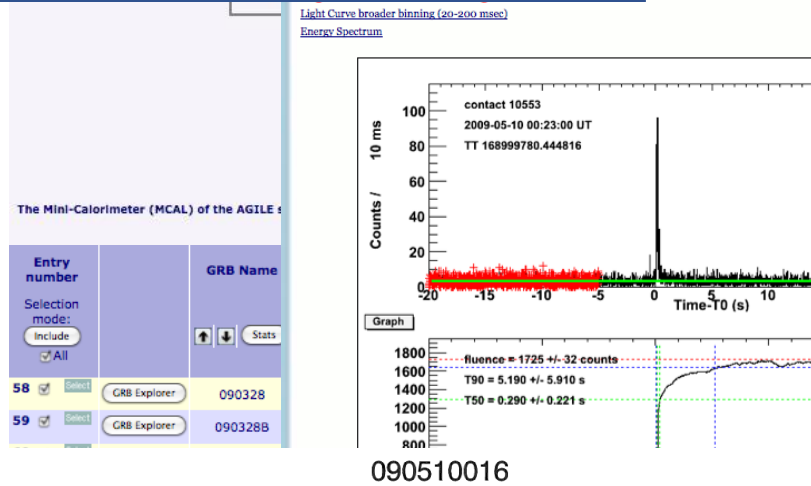
[Rebin this light curve](#) | [About these products.](#)

Flux Light Curve

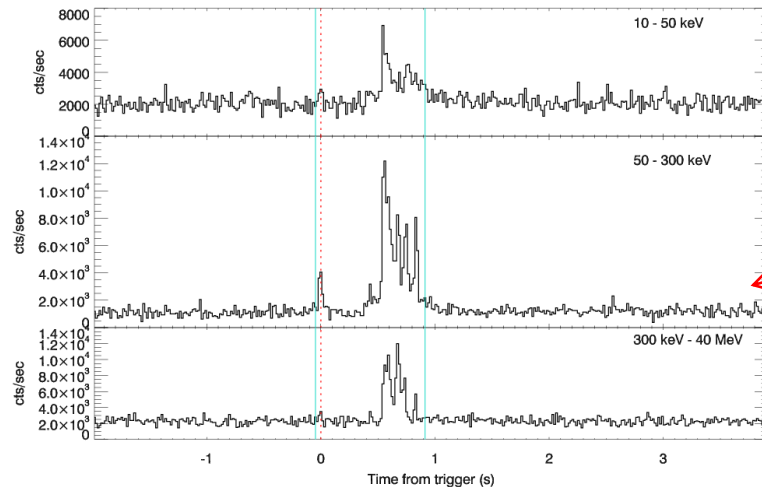
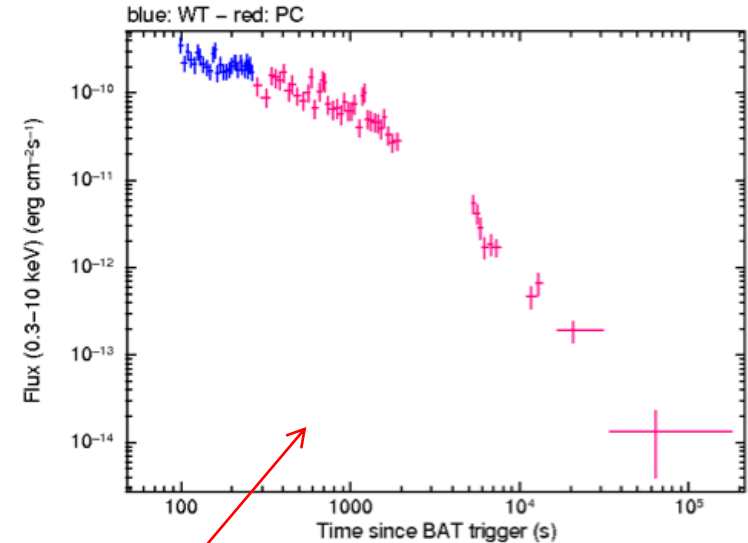
For this burst, 1 count = 4.0×10^{-11} erg cm^{-2} (observed flux) ([Automatic spectrum](#)).

Note that this is an average conversion factor: the true value may evolve with time.

[Rescale fluxed light curve.](#)



Swift/XRT data of GRB 090510



- Products
 - Swift-XRT light curve repository at Leicester
 - Swift-BAT
 - Quicklook GBM lightcurve
- GCN
- Blog for Gamma Ray Bursts
- Articles
- SAO/NASA Astrophysics Data System

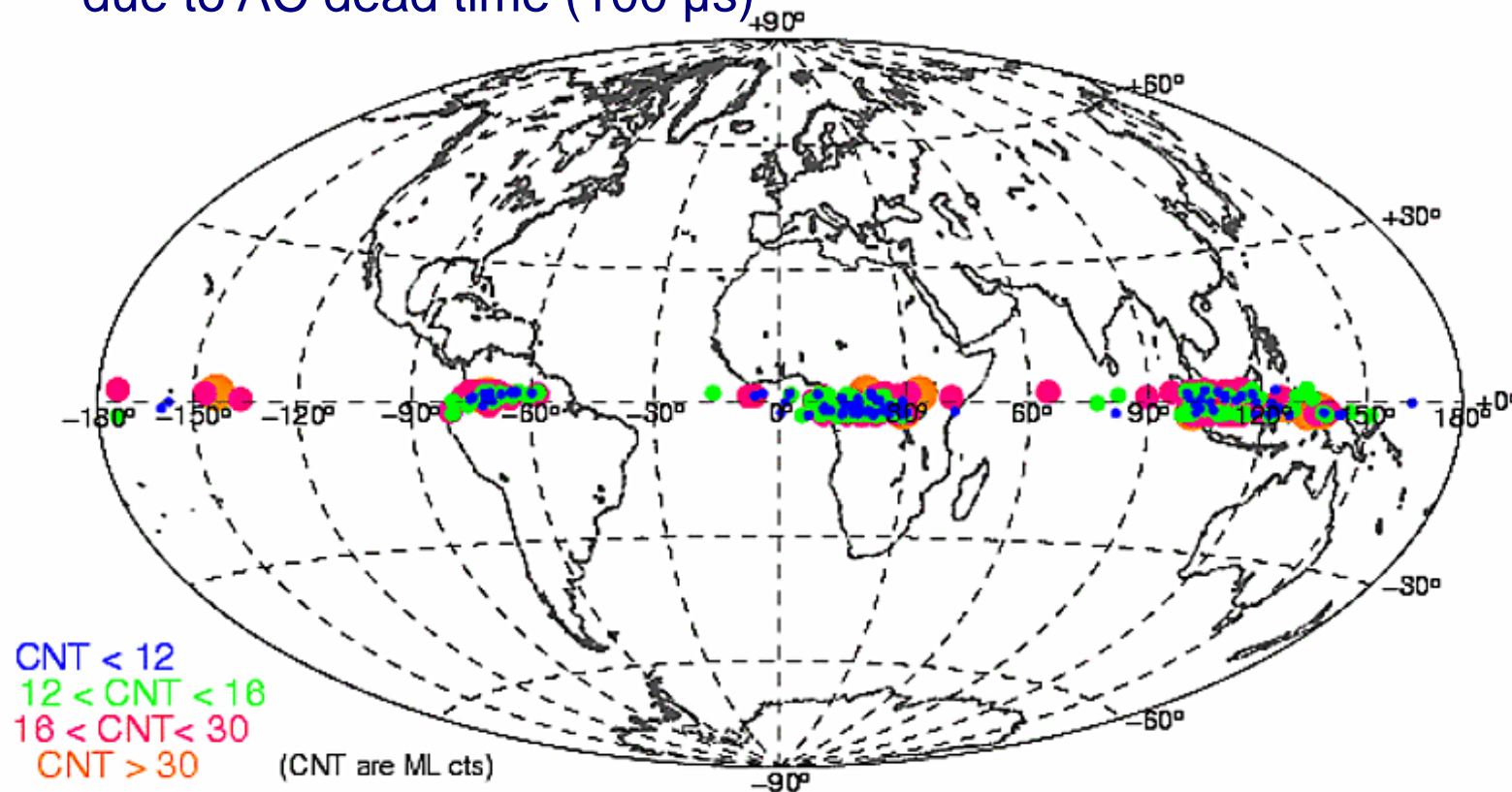
MCAL Terrestrial Gamma-ray Flash Catalogs

First catalog of MCAL TGF published in Marisaldi et al. 2014

“Properties of Terrestrial Gamma-Ray Flashes detected by AGILE MCAL below 30 MeV”
on data from March 2009 to July 2012

->detected 308 TGF below 30 MeV

- Longitudinal distribution similar to previous ones; comparison with RHESSI and Fermi/GBM sample; correlation with WWLLN radio on ground →larger durations due to AC dead time (100 μ s)



Catalog web
page to access
light curves

and new
dedicated tool

MCAL Terrestrial Gamma-ray Flash Catalogs

2nd catalog of MCAL TGF published in Marisaldi et al. 2015

“Enhanced detection of Terrestrial Gamma-Ray Flashes by AGILE”

on data from 23 March to 24 June 2015 ONLY with the AC Veto disabled!

->detected 279 TGF below 30 MeV

- TGF rate increased by 1 order of mag! Most with duration $< 100 \mu\text{s}$ as expected
- First unbiased by dead time sample \rightarrow correlation with Fermi/GBM sample
- Better correlation with WWLLN

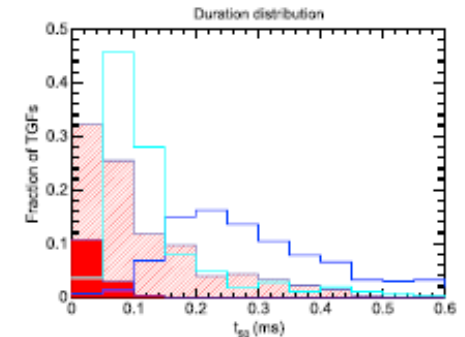
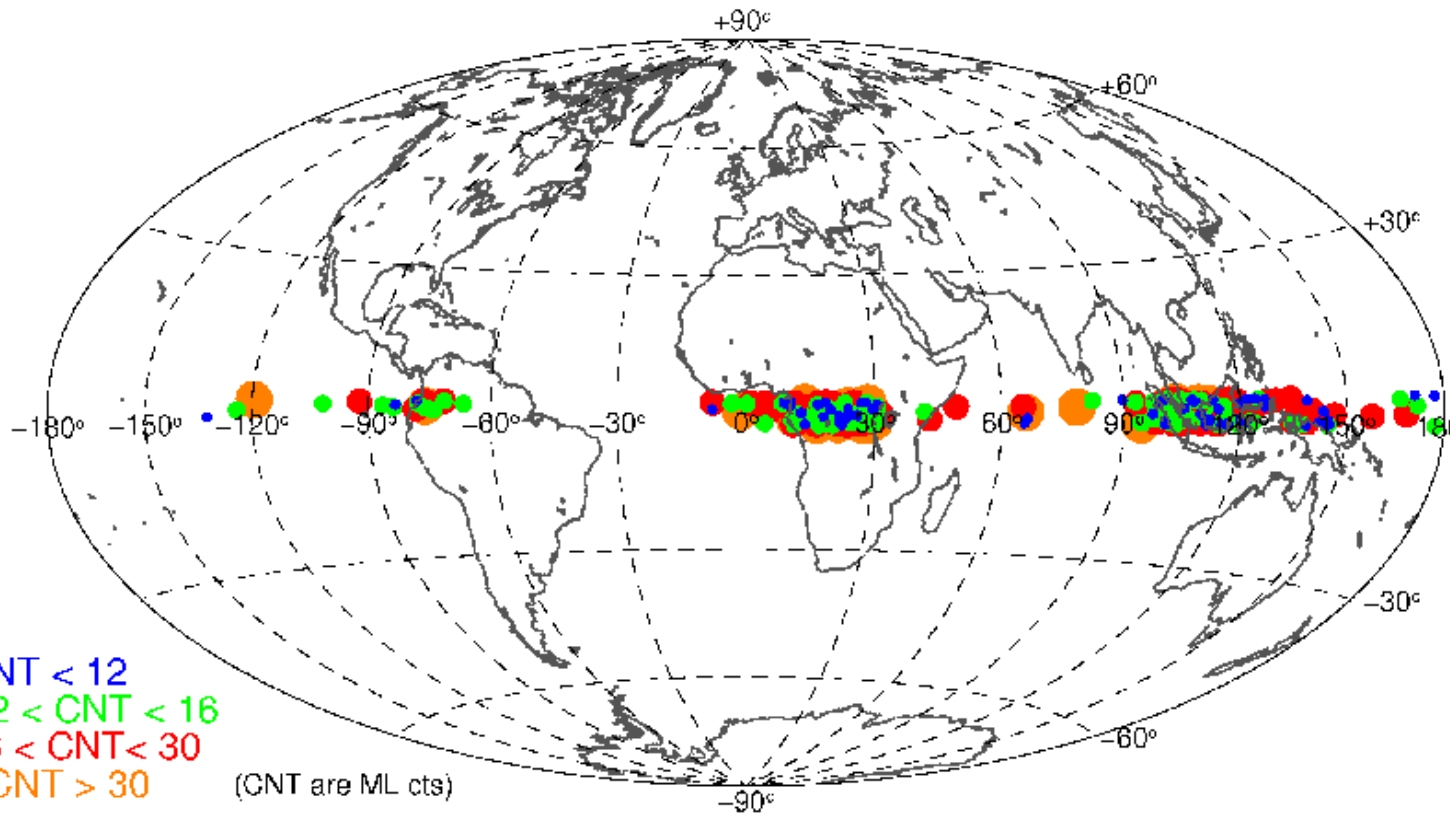


Figure 2. Normalized duration (t_{50}) distribution for the enhanced (red hatches) and standard (blue line) TGF samples. Red filled histogram: t_{50} distribution of the enhanced TGFs with a simultaneous WWLLN match normalized to the total number of the enhanced sample. Cyan histogram: Fermi GBM t_{50} distribution calculated for counts above 300 keV, from Connaughton et al. [2013], Figure 3.

MCAL Terrestrial Gamma-ray Flash Catalogs

Enhanced detection of Terrestrial Gamma-Ray Flashes by AGILE

TGF ($E < 30$ MeV) observed from March 2015 to June 2015

Help

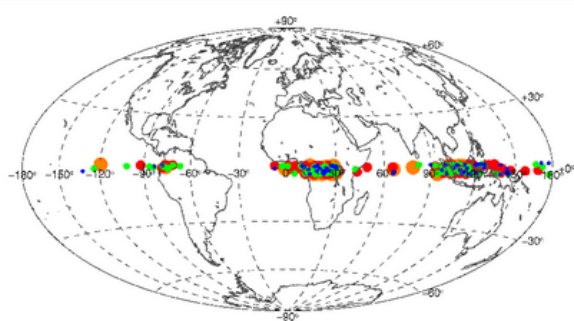
Show/hide columns

Advanced filtering

Print current view of table

Print complete table

Reset all filters



Search table columns

M. Marisaldi et al. 2015, Geophys. Res. Lett., 42, 94819487 (2015). Doi: 10.1002/2015GL066100 - Online version of the second AGILE Terrestrial Gamma-ray Flashes (TGF) catalog below 30 MeV detected by the Minicalorimeter (MCAL) instrument at configuration change aimed at the inhibition of the anti-coincidence veto signal on the MCAL instrument, in order to improve the TGF detection performance with respect to the previous configuration. The interactive web table includes 279 TGF detected by AGILE MCAL selected to have the maximum photon energy up to 30 MeV, and

Columns description

The first AGILE/MCAL TGF catalog on-line web table, Marisaldi et al. 2015

Export Current view of Table in:

Previous Page Next Page Page Size (# of lines) 200 Reset all filters Show all entries

This view includes 279 entries

Entry number	TGF ID	GeoLon	GeoLat	Date (UTC)	Trigger Time T0 (MET in s)	T0_micro (μ s)	T50 (ms)
1	TGF LC 150323.44645	-1.87	0.52	2015-03-23T10:42:54	354192174	410073	86
2	TGF LC 150323.70148	-132.93	-1.04	2015-03-23T16:50:08	354214208	140446	204
3	TGF LC 150323.85364	-71.92	2.46	2015-03-23T20:29:15	354227355	550377	61
4	TGF LC 150323.96107	119.48	-1.48	2015-03-23T23:03:57	354236637	71008	27
5	TGF LC 150324.21907	3.68	-2.12	2015-03-24T05:15:28	354258928	267858	40
6	TGF LC 150324.42976	5.27	0.92	2015-03-24T10:18:52	354277132	262591	103
7	TGF LC 150324.45177	118.06	1.5	2015-03-24T10:50:33	354279033	734131	67
8	TGF LC 150324.57062	8.24	2.4	2015-03-24T13:41:42	354289302	175046	86

The SSDC TGF Data Explorer - Mozilla Firefox

www.ssdc.asi.it/explorer_TGF.php 50%



Entry 150323
GeoLong. = 119.48
GeoLat. = -1.48

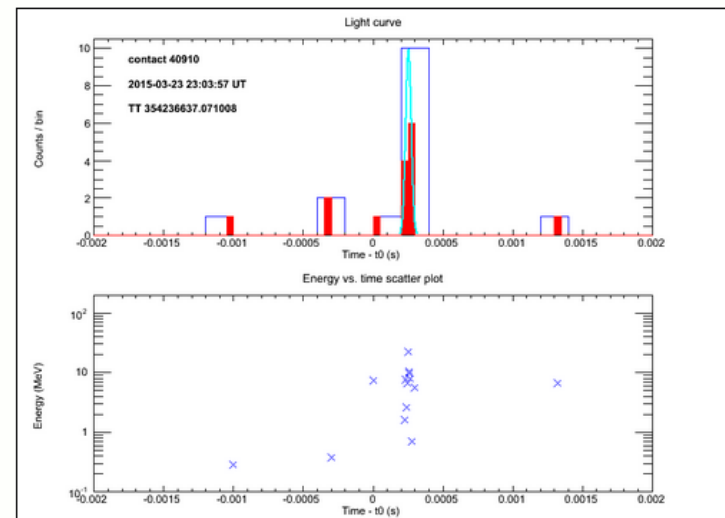


AGILE MCAL Data Products Source Details

Standard Products

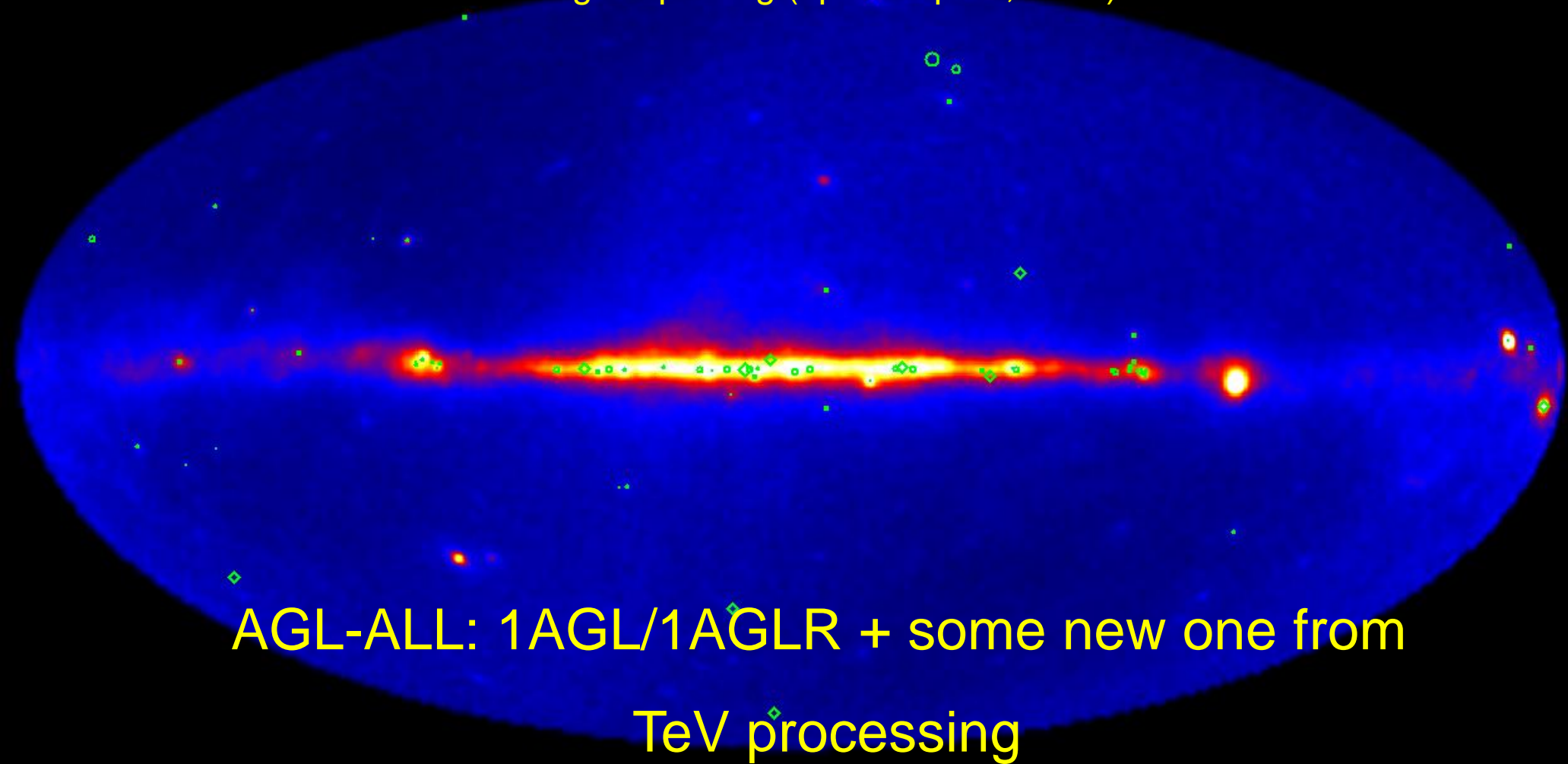
Light Curve Legend:

Blue histogram: 200 microsec time bin
Red filled histogram: finer binning 50 microsec
Cyan curve: maximum likelihood Gaussian fit



AGILE Total Intensity Map ($E > 100$ MeV):

Pointing + Spinning (up to Sep 30, 2017)



AGILE Total Intensity Map ($E > 100$ MeV):

Pointing + Spinning (up to Sep 30, 2017)

Short summary:

very soon news from GRID

but more still to discover in Spinning data archive

News will follow soon about MCAL “short” time scale events

2AGL

Thank you!

